

A New Insight into Newton' s Method for Optimization on Riemannian Manifolds

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Abstract

Newton' s method, a cornerstone algorithm for root-finding in Euclidean spaces, has been extended to Riemannian manifolds to tackle an increasingly broad class of optimization problems. While its computational characteristics and convergence properties have been extensively studied in the literature, the underlying algorithmic behavior of Riemannian Newton' s method remains inadequately characterized by existing theoretical frameworks—despite its proven practical effectiveness. This paper presents a novel theoretical perspective grounded in the principle of form invariance. We establish that the second-order Taylor expansion of a real-valued function possesses form invariance in local coordinates on a Riemannian manifold, which naturally gives rise to a fully quadratic model for Newton' s method. The resultant Newton step formula is identical to its Euclidean counterpart, preserving computational simplicity while accommodating the intrinsic curvature of the manifold. This work not only provides a rigorous explanation for the algorithm' s behavior but also paves the way for more efficient computational implementations.

Full Text

Preamble

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Abstract

Newton' s method, a cornerstone algorithm for root-finding in Euclidean spaces, has been extended to Riemannian manifolds to tackle an increasingly broad class of optimization problems. While its computational characteristics and conver-

gence properties have been extensively studied in the literature, the underlying algorithmic behavior of Riemannian Newton's method remains inadequately characterized by existing theoretical frameworks—despite its proven practical effectiveness. This paper presents a novel theoretical perspective grounded in the principle of form invariance. We establish that the second-order Taylor expansion of a real-valued function possesses form invariance in local coordinates on a Riemannian manifold, which naturally gives rise to a fully quadratic model for Newton's method. The resultant Newton step formula is identical to its Euclidean counterpart, preserving computational simplicity while accommodating the intrinsic curvature of the manifold. This work not only provides a rigorous explanation for the algorithm's behavior but also paves the way for more efficient computational implementations.

Keywords: Newton's method, Riemannian manifold, form invariance, optimization, second-Order Taylor Expansion MSC Classification: 49M15 , 65K10 , 53B20 , 41A58

1.1 Background

Optimizing a real-valued function on a Riemannian manifold has emerged as a central task in science and engineering, spanning a wide range of nonlinear constrained optimization problems [1–4]. In such problems, the feasible set constitutes a nonlinear space, rendering them substantially more challenging to solve than their unconstrained counterparts [5]. While traditional algebraic methods are applicable, differential geometry provides a powerful tool and natural language for characterizing these geometric properties, often facilitating more intuitive theoretical analysis and more effective numerical computation [6]. Riemannian optimization addresses this class of problems by treating the constraint set as a manifold. Algorithms within this framework perform a series of descent steps along geodesics, leveraging intrinsic geometry of the manifold.

This paper employs various distinct mathematical notations. To ensure clarity and consistency in notational conventions throughout the work, typical symbol usages are summarized in Appendix A.

Euclidean n -space \mathbb{R}^n , is the simplest example of a Riemannian manifold. Extending well-established optimization methods originally developed in \mathbb{R}^n to general Riemannian manifolds is a natural and widely adopted practice. The second-order Taylor expansion of an objective function f at a point $x \in \mathbb{R}^n$ with a perturbation $s \in \mathbb{R}^n$ takes a particularly concise form: $f(x + s) = f(x) + \partial f(x)^T s + \frac{1}{2} s^T \mathcal{H} f(x) s + O(\|s\|^3)$.

The core idea of Newton's method for minimizing f lies in iteratively applying Eq.1 to refine the current estimate along a descent direction. The steepest descent direction, $-\partial f(x)$, arises from a linear approximation of f , whereas the Newton step, $-\mathcal{H} f(x)^{-1} \partial f(x)$, stems from a quadratic approximation and represents the second-order steepest descent direction. Beyond its rapid

quadratic convergence rate, Newton's method frequently serves as a key component in other numerical algorithms, such as trust-region methods [5, 7].

Riemannian optimization fundamentally involves leveraging the local geometric information of the objective function and its constraints. However, due to the substantial differences in differential structures between a general manifold M and Euclidean space \mathbb{R}^n , it is widely acknowledged that the proper extension of Newton's method from \mathbb{R}^n to M necessitates a robust framework of differential geometry—including concepts like the Riemannian connection and covariant derivatives—as well as computationally intensive calculations of geodesics and parallel translation [1, 8, 9].

Furthermore, existing theory still lacks satisfactory explanations for certain algorithmic behaviors of Riemannian Newton's method. For instance, while local normal coordinates (induced by the exponential map) are employed to handle the Taylor expansion of f on a "curved manifold" [8], they do not adequately account for why Newton iterations retain robust quadratic convergence, even when the acceleration term is deliberately omitted [9]. The fundamental distinction between the Riemannian and Euclidean versions of Newton's method warrants further study. This paper explores these open questions from a new perspective.

1.2 Preliminaries

This section elaborates on the Taylor expansion of a differentiable function and its application to Riemannian optimization. Foundational concepts in Riemannian geometry can be found in [10-12].

1.2.1 Taylor expansion on Riemannian manifolds

Let M be a smooth manifold equipped with a Riemannian metric g . Consider a vector field X on M and its associated local flow ϕ_t for small $t \in (-\epsilon, \epsilon)$. For a point $p \in M$, the curve $c(t) = \phi_t(p)$ is the integral curve for X passing through p at $t = 0$. Let D be the set of functions on M that are differentiable at p . The tangent vector to this curve at $t = 0$ defines a linear operator $X_p = c'(0)$ on D , given by $d(f \circ c)|_{t=0}$. The collection of all such vectors to M at p forms the tangent space, denoted by $T_p M$.

The vector field $v(t) = c'(t) \in T_{c(t)}M$, is referred to as the velocity field of c .

The ordinary derivative of v along c is not intrinsically defined. To resolve this issue, we employ the covariant derivative D_{dt} , which is induced by the Riemannian connection ∇ . For a vector field Y induced by $X \in T M$ along c , i.e. $Y(t) = X \circ c(t)$, $dt Y = \nabla_{dc/dt} Y$. This enables us to define the intrinsic covariant derivative as D acceleration of the curve c as $a(t) = c''(t) = D_{dt} v(t) \in T_{c(t)}M$. The curve with zero acceleration is precisely a geodesic of M .

Under the Riemannian connection, the second-order Taylor expansion of f along

c is given by [9, 13]: $f \circ c(t) = f(p) + g_p(\text{grad}f, v) \cdot t + \frac{1}{2} g_p(\text{Hess}f(v), v) \cdot t^2 + O(t^3)$ (2) where $\text{Hess}f$ is the $(1, 1)$ -tensor known as the Riemannian Hessian of f .

In the literature, two common definitions of the Hessian exist. The present work adopts the $(0, 2)$ -tensor definition (See Appendix A), denoted as $\text{Hess}f$. These two Hessian tensors are naturally related by [11]:

$\text{Hess}f(X, Y) = g(\text{Hess}f(X), Y) = g(\text{Hess}f(Y), X)$ By expressing the directional derivatives of f in the directions of v and a , Eq.2 can be reformulated more compactly as $f \circ c(t) = f(p) + \nabla v f|_p \cdot t + \frac{1}{2} [\text{Hess}f(v, v) + \nabla a f]|_p \cdot t^2 + O(t^3)$.

The formulation using $\text{Hess}f$ is often preferred in optimization literature, particularly when M is a submanifold of a Euclidean space, since $\text{Hess}f$ can be computed by projecting the ambient Euclidean Hessian onto $T_p M$ (See Section 4.3).

$\text{Hess}f$ is positive definite if $\text{Hess}f(X, X) > 0$ for all non-zero vectors $X \in T_p M$.

Equivalently, $\text{Hess}f$ is positive definite if $g(\text{Hess}f(X), X) > 0$ for all non-zero X .

1.2.2 Optimality conditions

The optimality conditions in M are analogous to their classical counterparts in \mathbb{R}^n .

Specifically, a point $p \in M$ is defined as a critical point of f if its gradient vanishes at p —or equivalently, $\partial f = 0$ (or $df = 0$) [13, 14]. When the Hessian is non-degenerate, it provides crucial information about the nature of the critical point, as formalized in the following proposition:

Proposition 1 ([13]) Let $p \in M$ be a nondegenerate critical point of a smooth function $f: M \rightarrow \mathbb{R}$. Then p is a strict local minimum of f if and only if the Hessian (either $\text{Hess}f$ or Hf) is positive definite.

Pos.1 is not the most economical, as $\partial f = 0$ implies $\text{Hess}f = \text{Hf}$ (See Remark 1). The requirement of positive definiteness for $\text{Hess}f$ can be relaxed to that of its Euclidean counterpart Hf . Alternatively, a similar simplification can also be achieved through appropriate parametrization, even for non-critical points. The geodesic-induced exponential map centered at p gives rise to normal coordinates, where the Christoffel symbols vanish at p , significantly simplifying the local representation:

Lemma 2 In normal coordinates around p , $\text{Hess}f_p = \text{Hf}_p$.

While $\text{Hess}f$ generally depends on the metric, there is a noteworthy exception:

Lemma 3 (Lemma 68, [11]) If a function $f : M \rightarrow \mathbb{R}$ has a critical point at p , then $\text{Hess}f_p$ does not depend on the metric.

Remark 1 Strictly speaking, the expression $\text{Hess}f = \text{H}f$ constitutes a mild abuse of notation, as $\text{H}f$ is not a tensor of M . In local coordinates, the components of $\text{Hess}f$ are given by $\text{Hess}f(\partial_i, \partial_j) = \partial_i \partial_j f - \Gamma^k_{ij} \partial_k f$ [11], where Einstein summation notation is adopted.

$\text{Hess}f(\partial_i, \partial_j)$ reduce to $\partial_i \partial_j f$ (the components of $\text{H}f$) when either the Christoffel symbols vanish or at a critical point. It is in this sense that we use the notation $\text{Hess}f = \text{H}f$. Consequently, the Hessian is often identified with its matrix representation in a given coordinate system, and the two are used interchangeably.

Lemma 2 and 3 establish that in a normal coordinate system, the Taylor expansion in Eq.4 can be simplified to a form analogous to that in Eq.1. We now generalize this result, proving that—with appropriate technical treatments—this property is not unique to normal coordinates but extends to any admissible coordinate system.

1.2.3 Contributions

The preceding observations motivate an investigation of Riemannian Newton's method through the lens of invariance. The primary contributions of this work are as follows: (cid:136) We demonstrate that the second-order Taylor expansion of a differential function on a manifold exhibits form invariance under the Riemannian connection. (cid:136) Leveraging this invariant form, we formulate a purely second-order Riemannian Newton step. Crucially, it retains the same expression as the standard Euclidean Newton's method. (cid:136) This invariant formulation clarifies key characteristics of Riemannian Newton's method, naturally explaining its independence from the Riemannian connection and metric.

These properties summarized in Table 1 show that Riemannian Newton's method is locally equivalent to its classical Euclidean counterpart and provide insights for guiding the implementation and refinement of the method.

Table 1 Comparison of Descent Directions in Euclidean and Riemannian Newton's methods
 Manifolds Quadratic models 1st-order Descent Directions Steepest Standard gradf 2nd-order Descent Directions Steepest $\text{H}f - 1\partial f$ $\text{Hess}f - 1\partial f$ $\text{Hess}f - 1\text{grad}f$ This work $\text{G} - 1\partial f$ $\text{H}f - 1\partial f$ The remainder of this paper is organized as follows. Section 2 reviews the development of descent methods, with a particular focus on Riemannian Newton's method.

In Section 3, we utilize the Taylor expansion to analyze the invariance properties of Riemannian Newton's method, demonstrating that on a local chart, it is intrinsically equivalent to the standard Euclidean Newton's method. Section 4 compares the numerical characteristics and computational approaches of the existing Riemannian Newton's method with the one proposed in this work,

while Section 5 validates our findings with two illustrative examples. Finally, Section 6 concludes the paper.

2 Related work

Gradient descent is a fundamental parameter estimation method, renowned for its computational simplicity. A prominent instance is the steepest descent method; however, its convergence can be notably slow on ill-conditioned problems, as its iterates follow a zigzagging trajectory toward the solution. The conjugate gradient method offers accelerated convergence without requiring matrix storage. Since both methods use only the gradient of the objective function to determine search directions, they are classified as first-order methods. These approaches have since been extended to Riemannian manifolds, where numerous variants have been proposed and are widely applied [3, 4, 8, 13]. In contrast, Newton's method is a second-order technique that incorporates the Hessian. The resulting Newton direction is widely regarded as one of the most effective search directions available.

Historically, Newton's method is also referred to as the Newton-Raphson method.

Its original conception traces to François Viète, Isaac Newton, and Joseph Raphson in the 17th century, who laid the foundational idea for computing roots of polynomial equations. In 1740, Thomas Simpson adapted the method to nonlinear equations into a form closely aligned with its modern formulation, and further extended its applicability to minimization problems by setting the gradient to zero [15, 16]. The formal generalization of Newton's method to manifolds, by contrast, dates to the 1970s. It was not until Luenberger's seminal 1972 work that Newton's method was explicitly formulated and analyzed for constrained optimization problems via the Riemannian structure of constraint sets [17]. Gabay (1982) subsequently extended Newton's method and other descent methods to manifold settings, though this work did not formally employ the notion of Riemannian Hessian [13]. Smith (1994) and Udriste (1994) formulated Newton's method for locating zeros of a 1-form on a Riemannian manifold [8, 14]. Edelman et al. (1998) utilized matrix algebraic structures to develop Newton and conjugate gradient algorithms on Stiefel or Grassmann manifolds [18]. In these manifold-based approaches, the exponential map is typically used to retract the Newton step vector from the tangent space back to the manifold. To circumvent the high computational cost of geodesics, Manton (2002) proposed using a simpler projection operator for local parameterization [19]. Adler et al. (2002) formalized a unified framework of retractions—smooth maps from the tangent bundle to the manifold to execute projections [20]. Retractions have become a fundamental component of Riemannian computation [1, 9]. Absil and Malick (2012) showed that the definition of a retraction is highly flexible: any operation that maps a tangent vector back to the manifold along an "admissible" direction suffices [21]. This implies that Newton's method implementation imposes relatively lenient requirements on retractions. For a

systematic treatment of Riemannian Newton's method, we refer the reader to [1, 9].

Implementing Riemannian Newton's method requires specialized considerations critical to its performance. A primary challenge is the lack of global convergence for the method, especially when initial guesses are poorly chosen. While Kantorovich's theorem provides verifiable conditions at the initial point that establish convergence of the iterative process and the existence and uniqueness of the solution [22], it offers no criterion for selecting a suitable initial point. To ensure convergence from remote starting points, modified variants incorporating regularization [23], damping techniques [24], the Armijo condition [6], or adaptive trust-region strategies [7] can guarantee both global convergence and superlinear local convergence under mild assumptions. A detailed discussion of these implementation aspects lies beyond the scope of this work.

For a comprehensive overview of various approaches to implementing and refining the Newton scheme, we refer the reader to [2, 5, 9, 23] and the references therein.

The core of this generalization lies in establishing the Newton equation on the manifold M , which can be approached in three main ways.

Path-based Optimization This method conceptualizes optimization as a stepwise search for a path c on M that minimizes f [9, 13]. A quadratic model of f is derived from the Taylor expansion of f at p in $T_p M$ to characterize the local variation of f along c . By constraining c to have zero initial acceleration (i.e., assume it is a geodesic or a second-order retraction curve [9]), we obtain the model: $m_p(v) = f(p) + g_p(\text{grad}f, v) + \frac{1}{2}g_p(\text{Hess}f(v), v)$. Minimizing Eq.5 yields the Newton equation $\text{Hess}f(X) = -\text{grad}f$, and the corresponding step is $X = -\text{Hess}f^{-1}\text{grad}f$. Leveraging the assumption $\text{grad}f_p \neq 0$ enables a similar simplification.

Vector Field Zero-Finding This approach reframes the problem of identifying a critical point as locating a zero of a vector field Y [1, 20, 22]. Applying Newton's method for root-finding to Y requires its Jacobian $J_p(X) = \nabla_X Y$. Since the Jacobian of $Y = \text{grad}f$ is the Riemannian Hessian, the Newton equation $J_p(X) = -Y_p$ again results in the step $X = -\text{Hess}f^{-1}\text{grad}f$.

One-Form Taylor Expansion This method derives the Newton step by means of a Taylor expansion of the differential 1-form ω [8, 14]. By an abuse of notation, we denote by $(\nabla\omega)_p$ two distinct yet closely related entities: the bilinear form induced by the covariant derivative of ω evaluated at p , and the corresponding homomorphism from $T_p M$ to its dual space $T_p^* M$ [8]. The Newton equation is then formulated as $(\nabla\omega)_p(\cdot, X) = (\nabla_X \omega)_p = -\omega_p$. For $\omega = df$, $\nabla_X \omega$ coincides with $\text{Hess}f(\cdot, X)$ [14], which yields the step $X = -\text{Hess}f^{-1}df$. In local coordinates, it reduces to $X = -\text{Hess}f^{-1}\partial f$.

In this work, we adopt the path-based optimization framework. It provides a more intuitive characterization of the function's variation, thereby facilitating

the in-depth analysis presented in subsequent sections.

3 Invariance of Newton's method over a Riemannian

manifold Suppose the pair (U, x) be a chart on M , where the parameterization $x : U \rightarrow x(U) \subset M$ is a homeomorphism of some domain $U \subset \mathbb{R}^n$ on a subset $x(U) \subset M$. The coordinates x^1, \dots, x^n of a point $x \in U$ serves as the local coordinates of the point $p = x(x) \in M$. In the same sense we term this the local coordinate system of M at p . We adopt the Einstein summation convention in this section.

3.1 Invariance of the second-order Taylor expansion

By expressing the function f and the integral curve c in x , we obtain $f \circ x(x) = f(x^1, \dots, x^n)$, $x = (x^1, \dots, x^n)^T \in U$, $x^{-1} \circ c(t) = (x^1(t), x^2(t), \dots, x^n(t))^T$ respectively. Here f denotes, by a common abuse of notation, the coordinate representation of f in x . Therefore, restricting f to c , we compute its directional derivative as follows: $c'(0)f = f(x^1(0), \dots, x^n(0))$ (cid:12) (cid:12) (cid:12) (cid:12) $t=0$ (cid:18) $(x^i)'$ ∂ (cid:19) Let $p = c(0)$. In the basis $\{\partial_i|_{t=0}\}$ of $T_p M$, we define $X^i(t) = \frac{d}{dt} x^i(t)$.

For notational simplicity, all subsequent evaluations are carried out at $t = 0$; for instance, $X^i(0)$ is denoted as X^i , and the tangent vector $c'(0)$ admits the representation $X = X^i \partial_i$. By invoking the standard identities for the covariant derivative ($\nabla_X f = X^i \partial_i f$), the Hessian ($\text{Hess} f(\partial_i, \partial_j) = \partial_i \partial_j f - \Gamma^k_{ij} \partial_k f$) and the acceleration $\dot{X}^j = \frac{d}{dt} X^j$ (cid:17) along c [11], Eq.4 can be written as (cid:16) $\frac{d}{dt} X^k$ (cid:16) $\frac{d}{dt} X^k + \Gamma^k_{ij} X^i X^j$ (cid:17) $f \circ c(t) = f(p) + X^i \partial_i f \cdot t$ (cid:20) $X^i X^j (\partial_i \partial_j f - \Gamma^k_{ij} \partial_k f) + \frac{d}{dt} X^k \dot{X}^k$ (cid:19) (cid:21) $+ O(t^3) = f(p) + \frac{d}{dt} X^i \partial_i f \cdot t + \frac{d}{dt} X^j \dot{X}^j \cdot t + O(t^3)$ Denote $s(t) = x^{-1} \circ c(t)$. Expanding $s(t)$ using Taylor series around $t = 0$ in $T_p M$, we have $\Delta s = s(t) - s(0) = s'(0) \cdot t + \frac{1}{2} s''(0) \cdot t^2 + O(t^3)$, whence it follows that $X \cdot t = s'(0) \cdot t = \Delta s + O(t^2) = s'(0) \cdot t + \frac{1}{2} s''(0) \cdot t^2 = \Delta s + O(t^3)$ Without loss of generality, consider a parameterization of x centered at $p = x(0)$ such that $\Delta s = s(t)$. Given that $\Delta s = O(t)$ in a neighborhood of p , substituting the last two expressions into Eq.6 and collecting terms up to $O(t^3)$ yields: $f(s) = f(p) + \partial_x f(p)^T s + \frac{1}{2} s^T H_x f(p) s + O(|s|^3)$ Eq.7 establishes that the second-order Taylor expansion of $f \circ c(t)$ is form-invariant in coordinates upon extension from \mathbb{R}^n to M . We thereby get a more rigorous quadratic model function:

$F_p(s) = f(p) + \partial f(p)^T s + \frac{1}{2} s^T H_f p s$ The first- and second-order steepest descent directions are as follows: (cid:136) Steepest Descent Direction: $X^s d = -(G^{-1} \partial f)^T \partial x$. This is the well-known local coordinate representation of $\text{grad} f$. It is an invariant expression that also depends on the choice of metric. (cid:136) Newton Step: $N_n s = -(H_f - \partial^2 f)^T \partial x$. It is form-invariant that is independent of the choice of metric. This result implies that Riemannian Newton's method is locally equivalent to the classical Euclidean Newton's method.

The aforementioned results are summarized in 1.

3.2 Local convergence

Given a chart (U, x) , a family of local charts can be generated by defining $x_p(s) = x(x^{-1}(p) + s) = x(x + s)$ around $p = x(x)$. The parameterization x_p is now a local diffeomorphism centered at $p = x(x)$ and the differential $dx_p : T_x U \rightarrow T_p M$ is an isomorphism. If $\{e_i\}$ is the basis for $T_x U$, the differential acts as $dx_p(x_i e_i) = x_i \partial_i$ or, for the inverse, $dx^{-1}_p(X_i \partial_i) = X_i e_i$.

The retraction can be defined as $\phi_p = x_p \circ dx^{-1}_p$. Let $x_0 = x^{-1}(p_0)$ for the initial iterate p_0 . The iterative scheme is then given by: $x_{k+1} = x^{-1}(\phi_{p_k}(-(\text{Hf} - 1\partial f)^T \partial x)) = x^{-1}(p_k) - dx^{-1}((\text{Hf} - 1\partial f)^T \partial x)$ where $p_k = x(x_k)$.

By the definition that $\partial_i f(p) = \text{Di}(f \circ x)(x^{-1}(p)) = \partial \partial_{x_i} f(x_1, \dots, x_n)$ [12], Eq. 9 is in fact equivalent to the classical Newton-Raphson iteration in \mathbb{R}^n : $x_{k+1} = x_k - \text{Hf}^{-1} \partial f(x_k)$. Suppose that Hf is Lipschitz continuous in a neighborhood of a critical point p^* , and that Hf_{p^*} is positive definite. Then provided p_0 is sufficiently close to p^* , the sequence of iterates generated by Eq.9 converges quadratically to p^* . This is a straightforward consequence of the convergence properties of the classical Newton's method [5].

4 Comparative Analysis

This section presents a systematic comparison of the properties and computational challenges associated with Newton's methods derived from the two model functions Eqs.5 and 8.

4.1 The Distinct Roles of $\text{grad} f|_0$ and $c''(0) = 0$

Either of these two conditions reduces Eq.2 to Eq.5, yet their implications differ fundamentally in the derivation of Newton steps on Riemannian manifolds.

A retraction satisfying $c''(0) = 0$ is designated as a second-order retraction [9], with the exponential map being the canonical example. Notably, such a retraction still generates Newton iterates with a quadratic convergence rate, a property invariant to the choice of retraction [1, 9]. This feature can be verified directly via 7: the quadratic model is fully determined by the Euclidean gradient and Hessian, and is entirely independent of the specific form of the curve c . This naturally raises a key question: if the assumption $c''(0) = 0$ does not affect the quadratic convergence of the method, why does the formal derivation appear to rely on it? In particular, what rigorous mathematical procedure justifies the transformation from Eq.5 into Eq.8?

The answer lies in the approximation error. Denote the Riemannian-Euclidean $\text{ij}\partial_k f(\text{cid:3}) = \text{Hess} f - \text{Hf}$ and it is straightforward to verify Hessian difference as $Q(\text{cid:2}) \Gamma_k$ that $Q = O(\partial f)$ as $p \rightarrow p^*$. This permits expressing the model

$mp(v)$ as: $mp(v) = f(p) + \partial f(p)^T v + Fp(v) + O(\partial f \cdot v^2)$ $v^T Hfp$
 $v + (p \rightarrow p^*) v^T Qp v$ Although the leading terms of Eqs. 7 and 10, namely
 Fp , are formally identical, Eq.7 is derived from a more rigorously principled
formulation. Specifically, the model $Fp(s)$ employs a secant vector $s = \Delta s$,
whereas $mp(v)$ relies on the tangent vector $v = t \cdot s'(0)$. This discrepancy in
vector selection underpins the divergent mathematical foundations of the two
expressions, yet no such distinction is enforced in practical computations.

Remark 2 The core reason that mp give rise to a quadratically convergent New-
ton's method lies in $gradf(p^*) = 0$. This enables the term associated with
 a to be treated as a higher-order infinitesimal when p is sufficiently close to
 p^* . Since $c''(0) = 0$ holds for any geodesic c , the exponential map is often
directly utilized in the literature to derive Eq.5. It worth noting, however, that
the exponential map serves solely to eliminate the a -related term and acts as a
projection operator; it is not a critical prerequisite for quadratic convergence.

To summarize, we formalize the following key result:

Proposition 4 Fp and mp approximate f with the following error bounds: $f(s)$
 $= Fp(s) + O(s^3) = mp(s) + O(Qs^2 + s^3)$.

Moreover, $Fp = mp$ hold if either of the following condition is satisfied: (cid:136)
 p is a critical point of f . (cid:136) Normal coordinates centered at p are adopted.

4.2 Comparison of Newton steps: $Hf - 1\partial f$ and $Hessf - 1gradf$

The subsequent two propositions elucidate the intrinsic connections between
these two steps.

Proposition 5 $Hessf - 1\partial f = Hessf - 1gradf$. Proof The identity in Eq. 3 holds
for all tangent vectors $X, Y \in Tp M$. In local coordinates, this relation implies
 $Hessf = G \cdot Hessf$. We thus derive the equality below:

$Hessf - 1\partial f = Hessf - 1G - 1\partial f = Hessf - 1gradf$. For the analysis presented
in this section, we impose the norm consistency condi- tion: for example, the
vector norm is taken as the p -norm, and the matrix norm is its induced matrix
norm. Under the assumptions that Hf is invertible and all relevant functions are
Lipschitz continuous in an open neighborhood of p^* , we establish the following
Proposition 6 $Hf - 1\partial f - Hessf - 1gradf = O(\partial f^2), p \rightarrow p^*$.

Proof Let $P = Hf$ and I the identity matrix. By applying the Sherman-Morrison-
Woodbury formula, we proceed with the derivation: $Hf^{-1} - Hessf^{-1} = P^{-1}$
 $- (P - Q)^{-1} I - (I - QP^{-1})^{-1}$ (cid:17) $= P^{-1}$ (cid:16) $= P^{-1}$ (cid:16) $=$
 $P^{-1}Q(I - P^{-1}Q)^{-1}P^{-1} I - (I + Q(I - P^{-1}Q)^{-1}P^{-1})$ (cid:17) As $\partial f \rightarrow$
 0 , we have $(I - P^{-1}Q)^{-1} \rightarrow I$. It follows that $P^{-1} - (P - Q)^{-1} = O(\partial f)$.

Consequently, $Hf - 1\partial f - Hessf - 1gradf = (Hf^{-1} - Hessf^{-1})\partial f = O(\partial f^2)$

Remark 3 While the steps $Hf - 1\partial f$ and $Hessf - 1gradf$ yield identical convergence
rates, a subtle distinction exists: when the iteration point p is far away from

p^* , the term Q cannot be neglected. In this regime, mp does not constitute a strict quadratic model of f , as further illustrated in Example 2 (Section 5).

4.3 Computing Gradients and Hessians

While a coordinate-free framework is theoretically favored for geometric analysis, practical numerical computations demand a parametric (or algebraic) representation.

The merit of this representation stems from its direct compatibility with numerical implementation and its capacity to streamline algorithmic derivations. A manifold can be parameterized either intrinsically (through internal coordinates) or extrinsically (through embedding in a Euclidean space). For the sake of clarity, this and subsequent sections dispense with the Einstein summation notation; vector components are explicitly denoted with subscripts instead of superscripts.

Projection Method The computation of $\text{Hess}f = \text{grad}^2 f$ poses a challenge as Riemannian connections generally lack closed-form expressions. However, if M is a submanifold of a Euclidean space E , its Riemannian connection coincides with the orthogonal projection onto the tangent space of the standard vector derivative of a smooth extension of the Riemannian gradient vector field in E .

Consider a Riemannian submanifold of $M \subset E$ and a smooth function $f : M \rightarrow \mathbb{R}$.

We introduce a local extension \hat{f} of f to E , along with the corresponding extension \hat{g} of $\text{grad} f$. Let Proj_p denote the orthogonal projector onto $T_p M$ at $p \in M$. The following two propositions, which underpin the subsequent computations, are invoked herein [1, 9]:

Proposition 7 $\text{grad} f = \text{Proj}(\text{grad} \hat{f})$. **Proposition 8** $\text{Hess}f(X) = \text{Proj}(D^2 \hat{g}(X))$ A cost function is typically formulated as a function of coordinates in E constrained to a submanifold. If both the function and constraint admit concise algebraic forms in terms of $x \in E$ (e.g., matrix operations), the projection method tends to yield computational advantages [1, 9, 18]. Otherwise, the parameterization method based on intrinsic coordinates is preferable. We next demonstrate this with a class of implicitly defined manifolds.

Parametrization Method A prevalent form of manifold constraint is given by the implicit equation $h(z) = 0$, where $h : E_n \rightarrow \mathbb{R}^m$ denotes a smooth map with component functions h_i for $i = 1, \dots, m$. We assume that 0 is a regular value of h , meaning the Jacobian of h at each $z \in h^{-1}(0)$ is of full rank m . Under this regularity condition, the set $M = \{z \in E_n : h(z) = 0\}$ is an embedded submanifold of E_n with dimension $n - m$.

By virtue of the implicit function theorem, the regularity of h guarantees the existence of a local parameterization in a neighborhood of a point p . Specifically, we can define it as $z : U \subset \mathbb{R}^{n-m} \rightarrow E_n$ wherein $\bar{x} \rightarrow z = (\bar{x}, \phi(\bar{x})) = (x,$

y): here, $\bar{x} \in \mathbb{R}^{n-m}$ denotes independent parameters, while $y = \phi(\bar{x}) \in \mathbb{R}^m$ represents dependent (cid:12) coordinates. This parameterization construction requires the Jacobian submatrix $\partial h(\bar{x}, y)$ to be non-singular. Let $\bar{f}(\bar{x}) = f(\bar{x}, y)$. Elementary algebraic manipulations yield the following expressions for the first and second derivatives of \bar{f} with respect to \bar{x} : $\nabla_{\bar{x}} \bar{f} = \nabla_{\bar{x}} f + K \nabla_{\bar{x}} y$ where $L = \nabla_{\bar{x}} h$, $K = -\partial h L^{-1}$. The matrix derivative notations employed herein are defined as follows: $\partial \text{vec}(X) = \text{vec}(X) \otimes I_n$, $F = \text{vec}(F)$, $x = (x_1, x_2, \dots, x_n)^T$, $X = [F_{11} \dots F_{1q} \dots F_{pq}]$. The $\text{vec}(\cdot)$ operator stacks the columns of matrix X into a vector: $\text{vec}(X) = (x_{11}, \dots, x_{k1}, x_{21}, \dots, x_{k2}, \dots, x_{1l}, \dots, x_{kl})^T$. The symbol \otimes denotes the Kronecker product. I_n denotes the n -by- n identity matrix.

5 Examples

To illustrate the preceding theoretical results, we present two specific examples, focusing on a comparative analysis of the advantages and disadvantages of the two approaches in practical computations. For notional simplicity, we omit the index p in the function.

Example 1: Unit Euclidean Sphere Consider the unit Euclidean sphere $S_n = \{x \in \mathbb{R}^{n+1} : \|x\| = 1\}$ and the objective function $f(x) = b^T x$ defined on S_n .

Parametrization Method S_n is implicitly defined by $h(x) = \|x\|^2 - 1$. We first establish a suitable local parameterization. Without loss of generality, consider the chart (U, x) where $U = \{\bar{x} = (\bar{x}_1, \dots, \bar{x}_n)^T \in \mathbb{R}^n : \bar{x}_1 > 0\}$. The parameterization $x : U \rightarrow S_n$ is given by $x = (\bar{x}, \sqrt{1 - \|\bar{x}\|^2})^T$, where $x_i = \bar{x}_i$ for $i = 1, \dots, n$ and $x_{n+1} = \sqrt{1 - \|\bar{x}\|^2}$. The metric induced on S_n by its embedding in \mathbb{R}^{n+1} is $g = \nabla_{\bar{x}}^T \nabla_{\bar{x}} h$. Since $\bar{f}(\bar{x}) = f(\bar{x}, y) = b^T x$, direct computation yields: $\nabla_{\bar{x}} \bar{f} = \nabla_{\bar{x}} f + K \nabla_{\bar{x}} y$. Consequently, the Newton step is $N_{ns} = -\nabla_{\bar{x}}^2 \bar{f}^{-1} \nabla_{\bar{x}} \bar{f}$ represented in the local coordinate system of S_n at x . $\nabla_{\bar{x}} \bar{f} = y^T G^{-1} \nabla_{\bar{x}} \bar{f} = y^T \text{grad} f$, Projection Method Alternatively, we can compute the Newton step by treating S_n as an embedded sub-manifold of \mathbb{R}^{n+1} . We apply the analytic expression of $\text{grad} f$ to the extension of $\text{grad} f$ from S_n to \mathbb{R}^{n+1} , yielding $\text{grad} f = \text{Proj}(\text{grad} \hat{f}) = \text{Proj}(b) = b - (b^T x)x$, $\text{Hess} f(v) = \text{Proj}(D(\text{grad} f)(v)) = -(b^T x)v$.

The Newton step is then $N_{ns} = -\text{Hess} f^{-1} \text{grad} f = 1$ coordinates of \mathbb{R}^{n+1} .

Notably, the two Newton directions are directionally consistent with the Riemannian gradient, differing only by a scaling factor.

For manifolds given in an explicit parametric form or admitting a natural cover

by a set of coordinate charts, the parameterization method exhibits superior computational efficiency and straightforward implementation. Partitioning the independent variables, however, is non-trivial: while certain problems possess an inherent structure that naturally facilitates variable partitioning, such guiding principles lack universal applicability, and arbitrary partitioning may lead to suboptimal computational performance [13].

Example 2: Paraboloid Consider the paraboloid M defined by the equation $z = x^2 + y^2$. Its implicit representation is given by the zero-level set of the function $h(x, y, z) = x^2 + y^2 - z = 0$. We define a cost function on M as $f(p) = f(x, y, z) = z$. This is a strictly convex function with a unique global minimum at the origin.

The x - y plane serves as a natural parameter domain for M . Consequently, computing the Newton step proves straightforward:

$$n_{ns} = -\frac{1}{2} \frac{\partial h}{\partial z} = \frac{1}{2} \begin{pmatrix} 2x \\ 2y \\ -1 \end{pmatrix} = \begin{pmatrix} x \\ y \\ -\frac{1}{2} \end{pmatrix}$$
 This update rule converges to the optimal point in just one iteration, representing the standard full second-order Newton step.

For comparison, we now compute the step using the projected Hessian Hess_f .

The unit normal vector to the surface is $n = \frac{\partial h}{\partial h}$, where $a = 4x^2 + 4y^2 + 1$. The Euclidean gradient of f is $\text{grad } f = (0, 0, 1)^T$. The Riemannian gradient on the manifold is: $\frac{\partial h}{\partial h} = (2x, 2y, -1)^T$ $\text{grad } f = \text{Proj}(\text{grad } f) = \text{grad } f - \langle \text{grad } f, n \rangle n = \begin{pmatrix} x \\ y \\ 2x^2 + 2y^2 \end{pmatrix}$.

Calculating Hess_f is more involved. Since M can be parameterized as $r(x, y) = (x, y, x^2 + y^2)^T$, the vectors of $\{r_x, r_y\}$ form a basis of $T_p M$ at the point $p = r(x, y)$.

Let $u = ur_x + vr_y$ be a vector in $T_p M$. Extending $\text{grad } f$ analytically from M to \mathbb{R}^3 , we have $D(\text{grad } f) = D(\text{grad } f) - D(\text{grad } f, n)n = \text{grad } f, n Dn$, and thus $\text{Hess}_f = -\text{Proj}(\text{grad } f, n Dn) = \text{Proj}(Dn)$. Applying this to u gives $\text{Proj}(Dn(u)) = a^{-1} \text{Proj}(V)$, $\text{Hess}_f(u) = u(4y^2 + 1) - 4vxy v(4x^2 + 1) - 4uxy$ where $V = \begin{pmatrix} 2x \\ 2y \\ -2 \end{pmatrix}$. It can be verified that $\text{Proj}(V) = V - \langle V, n \rangle n$ with $B = (4y^2 + 1)^2 + 16x^2y^2 - 8xy(2x^2 + 2y^2 + 1) - 8xy(2x^2 + 2y^2 + 1) (4x^2 + 1)^2 + 16x^2y^2$ $a^2 B^{-1}u$, and $\bar{u} = (u, v)^T$.

Thus, $\text{Hess}_f(u) = 2a^6 B^{-1}u$. The Newton equation in matrix form is $B^{-1}u = c, 2 \text{grad } f$. Note that the three scalar equations from $B^{-1}u = c$ are linearly independent where $c = -a^6 \text{grad } f$. A closed-form solution can be obtained using the pseudo-inverse of B :

$$n_{ns} = (B^T B)^{-1} B^T c = -\frac{1}{8} \begin{pmatrix} 2(1+2y^2) \\ 2(1+2x^2) \\ 2(x^2 + y^2) \end{pmatrix}$$

Comparing the results from Eq.12 and Eq.14 aligns with the expectations of Proposition 6.

Remark 4 $\text{Hess}_f - 1 \text{grad } f$ is parametrization-independent but requires computing

the Riemannian connection; whereas $Hf - 1\partial f$ is parametrization-dependent, does not involve the metric, and eliminates the need to compute the Riemannian connection. The choice between them can be made based on the specific characteristics of the application.

6 Conclusion

All n -dimensional Riemannian manifolds share the same infinitesimal structure [11].

Specifically, a manifold M in a neighborhood of a point p can be approximated by its tangent space $T_p M$. By analogy, a function f defined on M can be approximated via its Taylor expansion in $T_p M$. This paper investigates Riemannian Newton's method from the perspective of Taylor-expansion form invariance of f and demonstrates that the Euclidean and Riemannian Newton's methods—corresponding to quadratic models Eq.5 and Eq.8, respectively—exhibit strong mathematical consistency. Most crucially, Riemannian Newton's method is locally equivalent to its Euclidean counterpart.

Specifically, this work reveals several key invariance properties: (cid:136) The second-order Taylor expansion of f on M is form-invariant. (cid:136) The Newton step, $Hf - 1\partial f$, derived from this quadratic model, is itself form-invariant in coordinates, generalizing the Euclidean Newton step from R^n to (cid:136) The step $Hf - 1\partial f$ is independent of the Riemannian metric and connection. (cid:136) The step $Hf - 1\partial f$ is fully second-order.

These properties yield promising optimization strategies for Riemannian Newton-type algorithms. Since strict projection to constrain each iterate vector within a coordinate chart to the manifold is unnecessary—provided all iterates remain within the same chart until a successive chart update is required—a well-designed algorithm may thus avoid computationally expensive operations such as parallel translation.

Moreover, the first-order steepest descent direction is likewise independent of the Riemannian connection. These considerations confer substantial benefits for the design and implementation of manifold-based algorithms, as they reduce computational overhead while preserving theoretical rigor. For example, in trust-region methods, one must approximate $Hessf$ or $Hessf$ with a suitable symmetric matrix [1, 7, 9]. The results of this work demonstrate that Hf can be reliably utilized for algorithm design; this approach is not only simpler but also circumvents potential concerns regarding additional loss of accuracy.

Appendix A Some Notations We use the following notational conventions: scalars are italicized, vectors are in bold italic, and matrices or tensors are in boldface. Vectors are column vectors by default (e.g., $x = (x_1, x_2)^T$). Common symbols are summarized in Table A1, with local exceptions noted inline. When no ambiguity arises, the subscripts p or x are typically omitted. To streamline derivations, Section 3 employs the Einstein summation con-

vention with superscripted indices for vector components. In all other sections, this convention is discarded, with either coordinate-free notation or standard subscripted coordinate notation employed.

Table A1 Notation Summary Notation Description $H_x f$, $\text{grad}_p f$ (X, Y) $\text{grad}_p f$ $\text{Hess}_p f$ $\text{Hess}_p f$ n -dimensional Euclidean space endowed with an inner product. When n is clear from context, it is abbreviated to E .

Standard E_n equipped with its canonical inner product. k -th component of a vector $x \in E_n$. When Einstein summation is not used, the traditional notation x_k is retained.

Euclidean gradient of a scalar function f with respect to $x \in \mathbb{R}^n$, i.e., $\partial_x f = (\partial_1 f, \dots, \partial_n f)^T \in \mathbb{R}^n$. In the absence of ambiguity, abbreviated to $\partial f = (\partial_1 f, \dots, \partial_n f)^T$.

Classical Euclidean Hessian of f . Vector fields on a manifold M .

A parameterization of the bijective map $x : U \subset \mathbb{R}^n \rightarrow x(U) \subset M$, often termed a chart in the literature. For notational convenience and following standard convention, we denote the pair (U, x) as a chart, while reserving "parameterization" for x itself. Notably, these two notions are conceptually equivalent.

Coordinate vector fields on M , abbreviated to ∂_i for brevity.

Formal column vector formed by stacking basis vectors $\partial_1, \dots, \partial_n$, i.e., $\partial_x = (\partial_1, \dots, \partial_n)^T$. Riemannian metric on M , often denoted by the inner product symbol $\langle \cdot, \cdot \rangle$ too. Its matrix representation is G .

Value of g at point $p \in M$, evaluated on $X, Y \in T_p M$.

Riemannian gradient of f at $p \in M$, defined as the vector field satisfying $\text{grad}_p f(X) = \text{df}_p(X)$ for all $X \in T_p M$.

Exclusively denotes the Riemannian (Levi-Civita/canonical) connection in this work.

Covariant Hessian of f at $p \in M$, a symmetric $(0, 2)$ -tensor defined by the Lie derivative 1 Alternative definition of the Riemannian Hessian of f , a self-adjoint $(1, 1)$ -tensor given by $\text{Hess}_p f(X) = \nabla_X \text{grad}_p f$ for all $X \in T_p M$.

Christoffel symbols of the second kind.

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Note: Figure translations are in progress. See original paper for figures.

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